



Transport Studies in HSX at 1 Tesla



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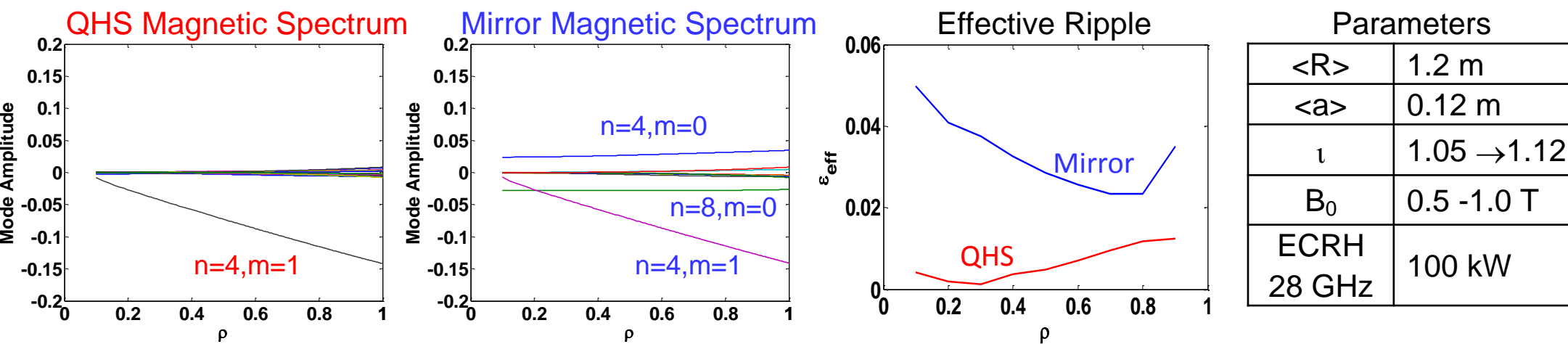
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Overview

- Thomson scattering profiles for several injected powers demonstrate that the core electron temperatures are 2-2.5x greater and that the density profiles are more peaked for the QHS as compared to Mirror for the same injected power.
- To match the QHS and Mirror profiles globally 2.3x the injected power (1.8x the absorbed power) is required for Mirror.
- The effective electron thermal diffusivity is 3-4x lower in the core for QHS when the profiles are matched.
- For a given temperature gradient the particle and heat fluxes in the core are lower for QHS.
- Several additional effects need to be taken into account to make comparisons between neoclassical and experimental values.

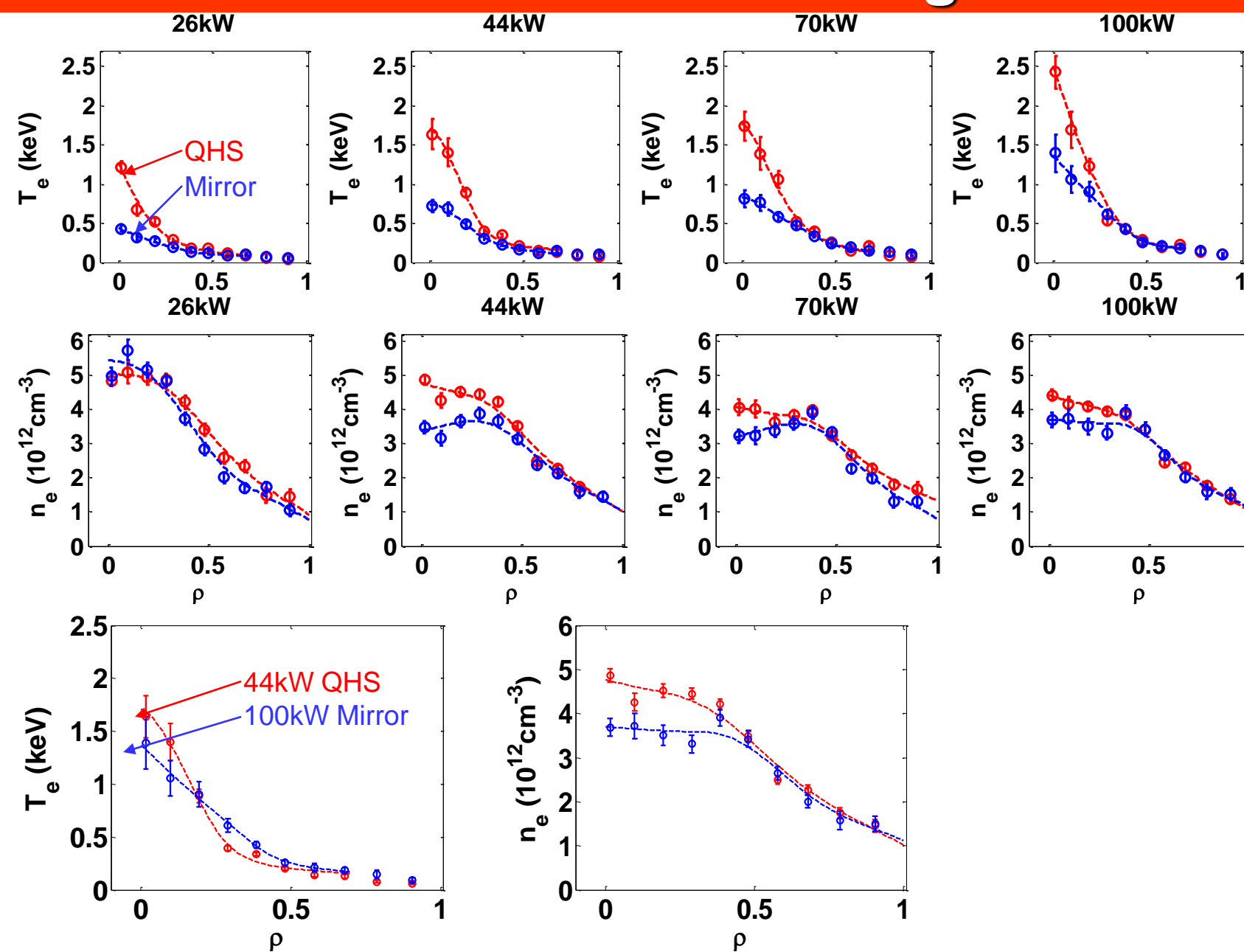
The HSX Stellarator



1. Profiles from Thomson Scattering

Power Scan Comparison

- For each injected power the core electron temperatures are 2-2.5x greater in QHS.
- Density profiles in QHS are more peaked than Mirror.
- The differences are greatest in the core, where neoclassical transport is important.
- The differences in these profiles indicate that the transport properties change between the two configurations.



Matching Profiles

- Due to the difference in the temperature gradients, the profiles cannot be exactly matched between the configurations.
- $\sum_n (T_n^{QHS} - T_n^{Mirror}) = 2eV$
- To minimize the difference between the profiles globally, 2.3x the injected power (1.8x the absorbed power) is needed for Mirror.

2. Particle Transport

Experimental Calculations

• The 3D neutral transport code DEGAS has been used to calculate neutral quantities including the particle source rate and H_α emission.

• The results are then coupled to measurements from a suite of absolutely calibrated H_α detectors to yield the source rate in absolute units.

• Due to the plasma densities ($n_e < 6 \cdot 10^{12} \text{ cm}^{-3}$), the source rate remains significant into the core.

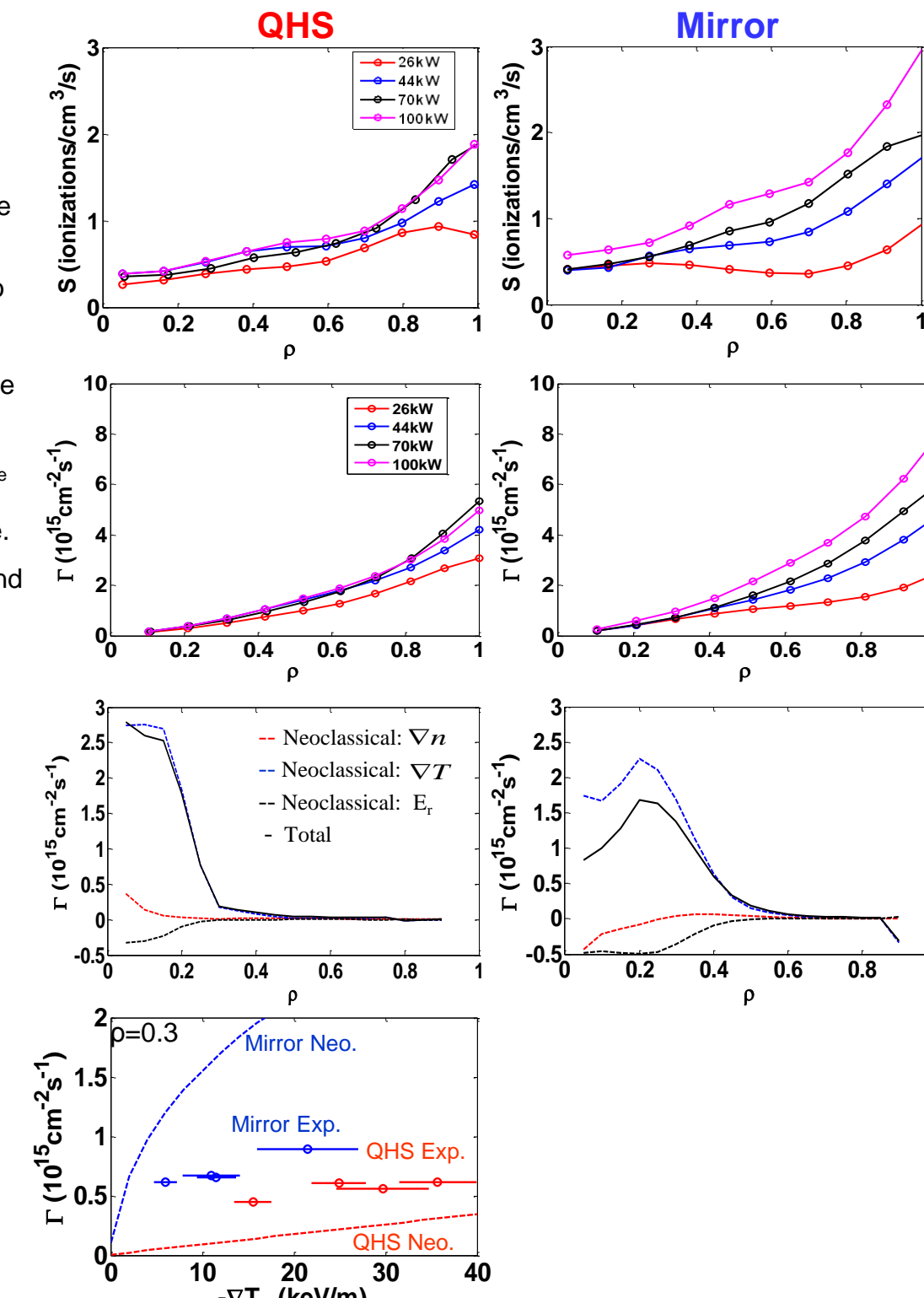
• Both the particle source rate and the particle flux scale with injected power in each configuration.

Neoclassical Calculations

• In each configuration the neoclassical particle flux in the core is dominated by thermodiffusion.

Core Particle Flux Scaling

• For a given temperature gradient the particle flux is larger in Mirror. This would result in a more hollow density profile.



3. Electron Thermal Transport

Heat Flux Calculation

• A ray-tracing code is used to calculate the power deposition profiles based on the local plasma density and temperature from the Thomson scattering profiles.

• The normalized power deposition profiles have the same width for each configuration for the Thomson profiles of interest.

• The power deposition profiles are volume integrated and scaled to the absorbed power measured from the diamagnetic loop.

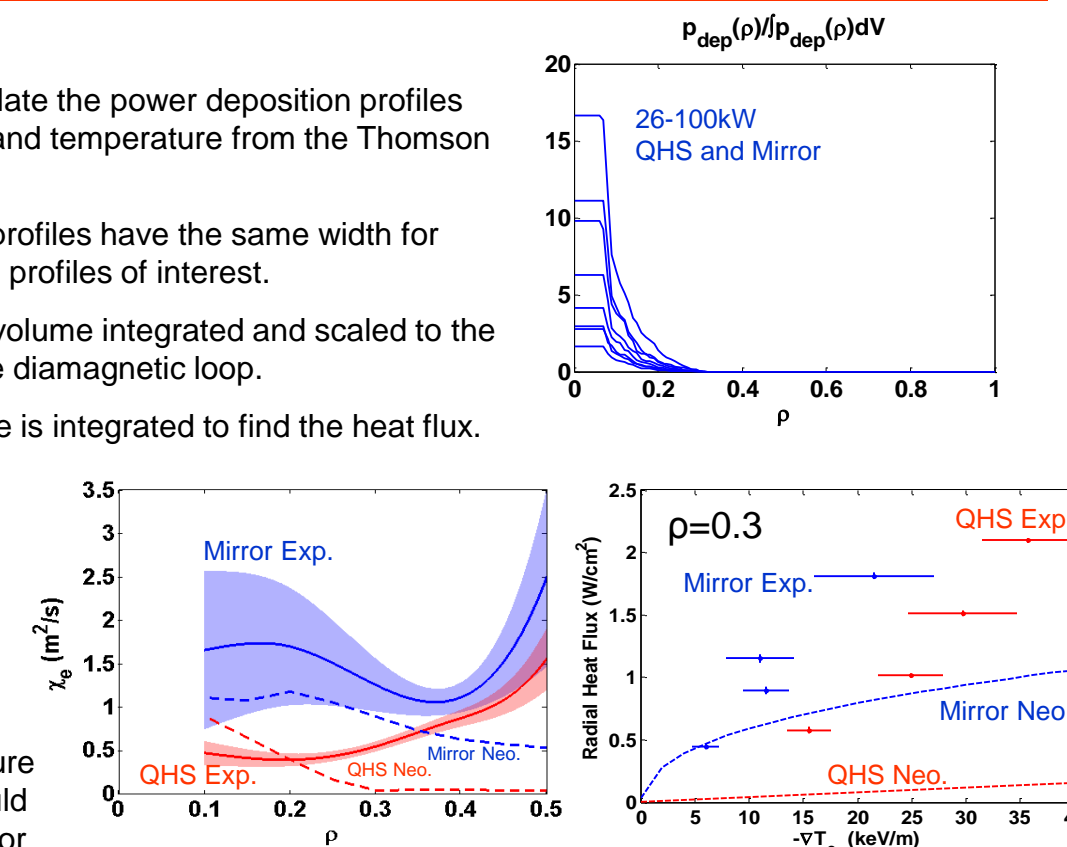
• The scaled power deposition profile is integrated to find the heat flux.

Effective Electron Thermal Diffusivity

• For the matched profile case χ_{eff} is lower in the core for the QHS configuration.

Core Heat Flux Scaling

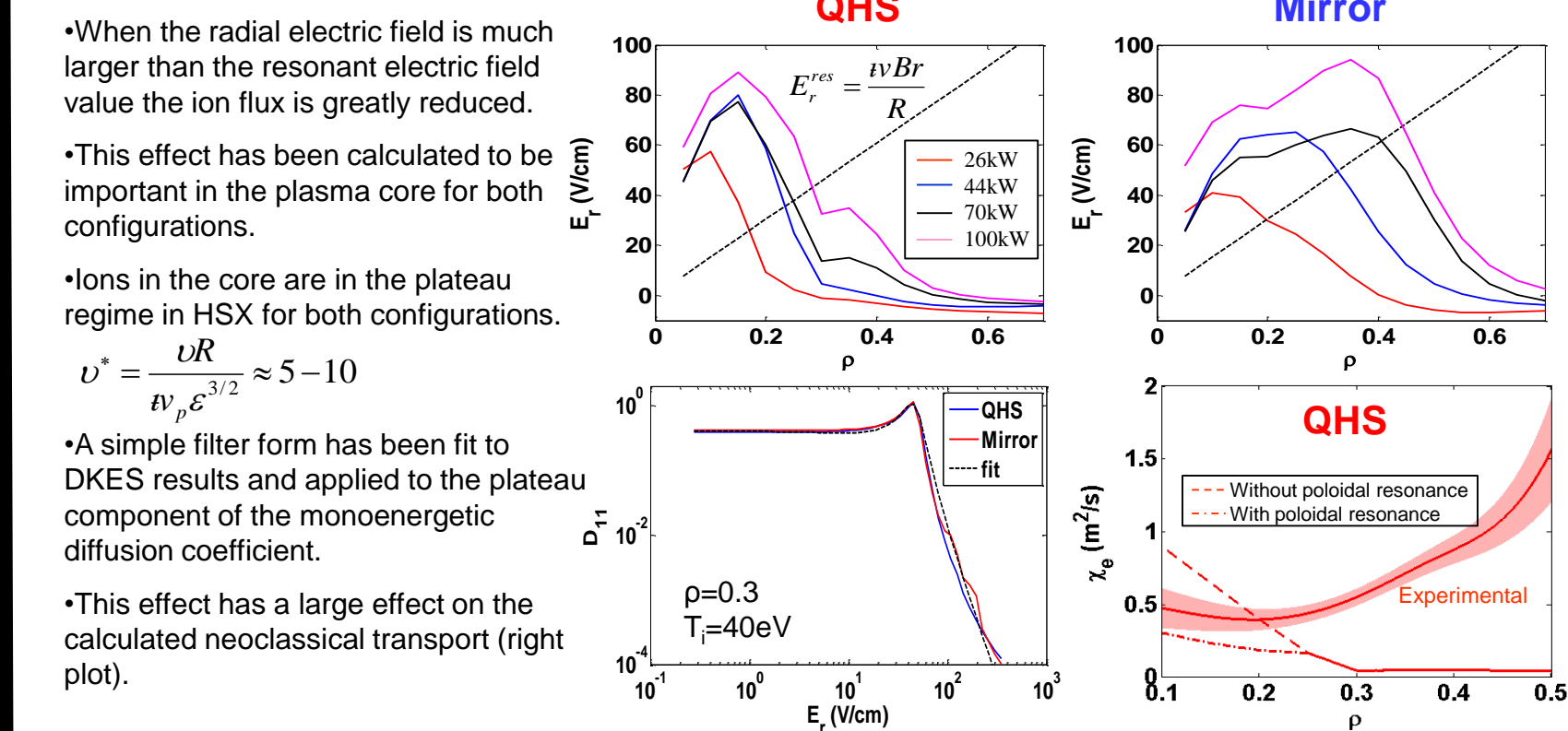
• For a given heat flux the temperature gradient is larger in QHS. This would result in a larger core temperature for similar edge conditions.



4. Neoclassical Calculations

Future Improvements

- For several cases the neoclassical calculations result in values larger than the experimental values.
- Several additional effects can be taken into account to reduce this discrepancy.
 - 1) Proper accounting of flows and momentum conservation using PENTA (Spong, 2005).
 - 2) The effect of the poloidal resonance on ions in the plateau regime (Maassberg *et al.*, 1993).



- 3) The ECRH driven flux can be calculated using the 5-D Fokker-Planck code GNET (Murakami, 2000).

• Results at B=0.5 Tesla (2nd Harmonic X-mode heating) indicate that this is an important effect in the core.

• The effect is predicted to be smaller at B=1.0 T (fundamental harmonic O-mode heating), but may still be significant.

- 4) Drift kinetics may not be valid near the core.

5. Confinement Scalings

Confinement time scaling for QHS

$$\tau^{exp} \sim n^{0.36 \pm 0.04} P^{-0.36 \pm 0.03}$$

N=244

RMSE=10.5%

• The scaling is weaker than ISS95/04

$$\text{Enhancement factor: } \frac{\tau_{QHS}}{\tau_{ISS04}} = 0.42$$

