# **Evidence for Fast-Electron-Driven Alfvénic Modes in the HSX Stellarator**

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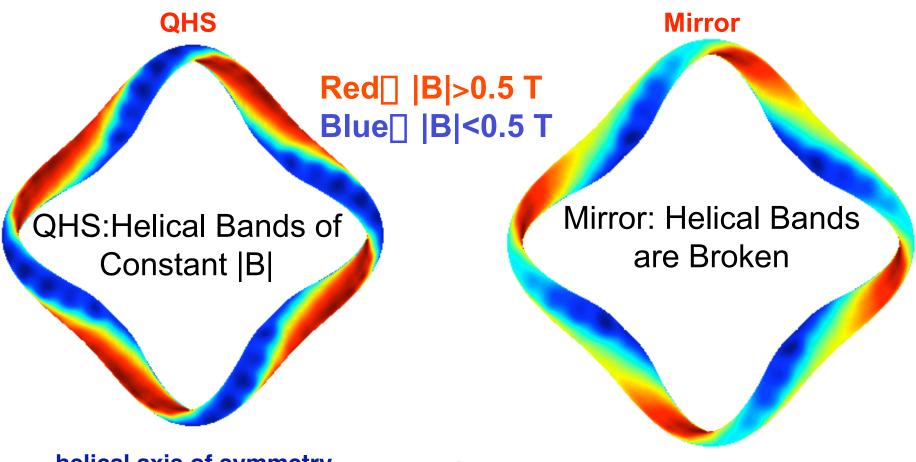






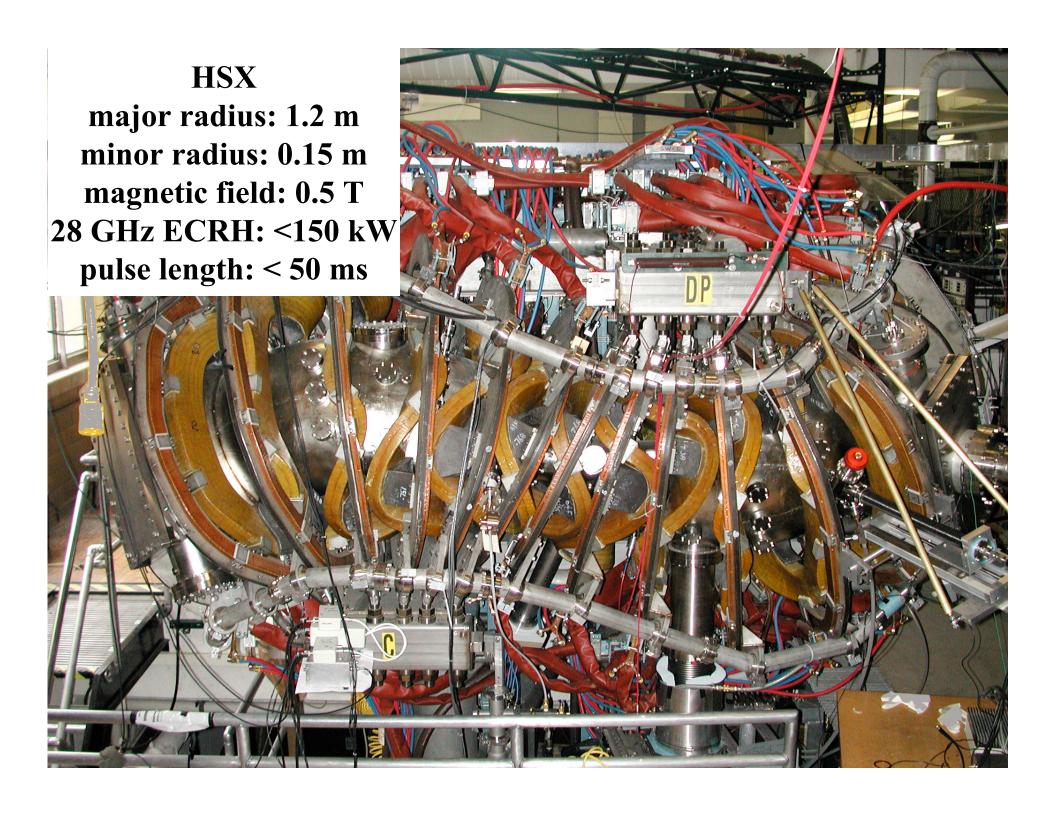
# HSX Provides Access to Configurations With and Without Symmetry

QHS: helical axis of symmetry in IBI; predicted very low neoclassical transport Mirror: quasi-helical symmetry broken by adding a mirror field.



helical axis of symmetry, no toroidal curvature, no toroidal ripple

Conventional stellarators exhibit poor neoclassical transport in low-collisionality regime due to magnetic field ripple



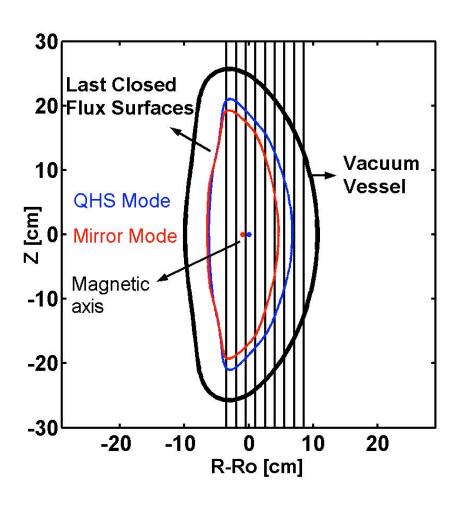
### **Outline**

- 1. Characteristics of observed fluctuations
  - Quasi-Helically Symmetric (QHS) configuration
  - Mirror (MM) configuration (conventional stellarator)
- 2. Alfvén Continua for QHS and Mirror Mode Plasmas (conventional stellarator) in HSX
- 3. Evidence for fast-electron driven GAE mode
- 4. Effect of biasing on Alfvenic mode

#### GOAL

- 1. Observe Alfvénic modes driven by fast electrons
- 2. Quasi-Helical Symmetry makes a difference

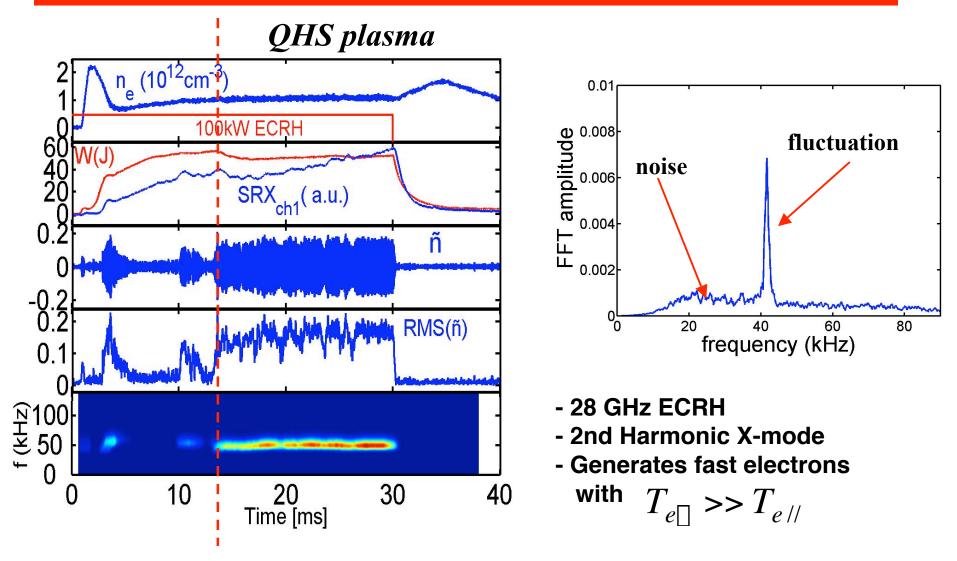
## Flux Surfaces and Interferometer Chords



#### Interferometer System:

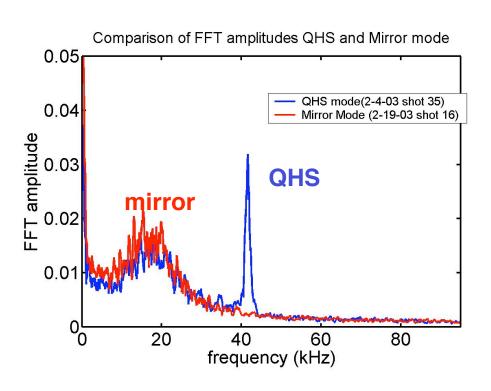
- 1. 9 chords
- 2. 200 kHz B.W.
- 3. 1.5 cm chord spacing

# **Coherent Density Fluctuations**

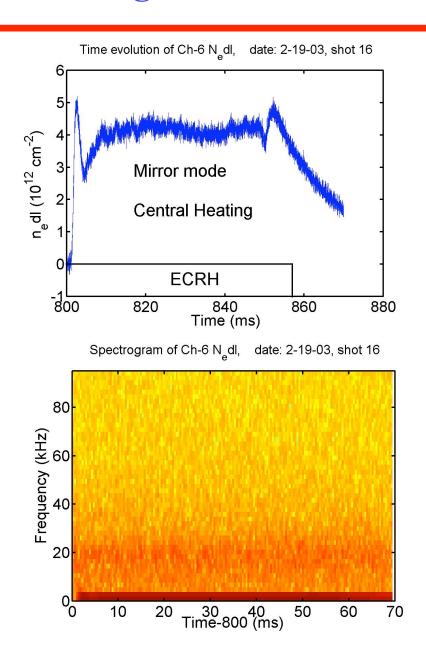


For  $P_{ECRH} > 100$  kW, confinement degrades Mode perturbs particle orbits leading to enhanced loss

## No mode observed in Mirror Configuration Plasma

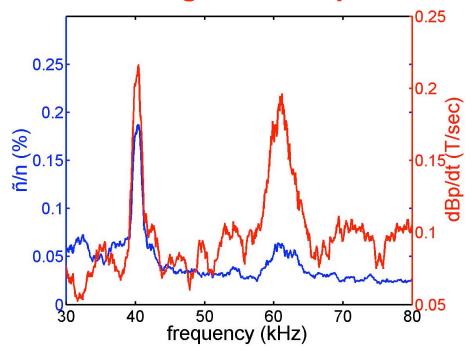


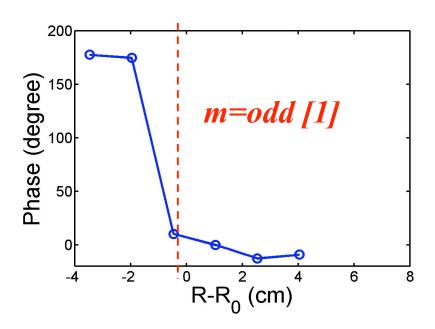
10% Mirror perturbation

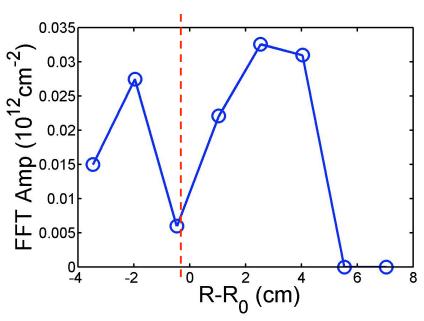


# **Fluctuation Features**

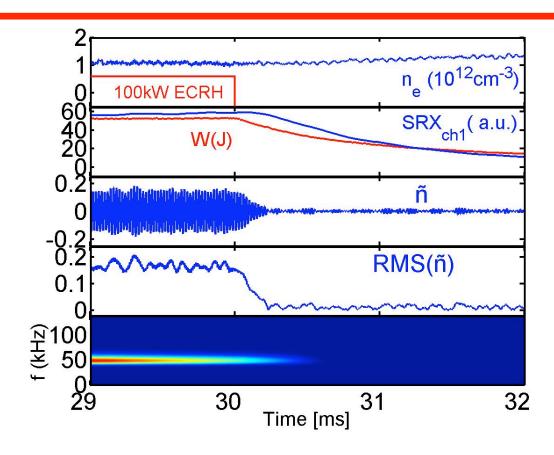
- only observed in QHS plasmas
- coherent, m=1 (n=?)
- localized to steep gradient region
- propagates electron drift direction
- Electromagnetic component







#### **Observed Fluctuations Associated with ECRH**



- Mode disappears ~0.2 msec after ECRH turn-off,
- faster timescale than W<sub>E</sub> and soft x-rays
- 2nd Harmonic X-mode generates nonthermal electrons (ECE) (no source for fast ions: T\_i~20 eV)  $T_{e \sqcap} >> T_{e //}$

Modes are driven by energetic electrons

#### Alfvénic Modes

Historically, Alfvénic modes have been observed on tokamaks or stellarators with NBI or ICRF to generate fast particles.

Alfvénic modes are generated if

1. resonance condition:  $V_p \ge V_A$  (V<sub>p</sub>: particle velocity) for trapped particles,  $\bigsqcup_{Dh} = \bigsqcup_{Alfven}$ 

where  $\square_{Dh}$  is the trapped-particle precessional drift frequency,

depends on particle energy, not mass

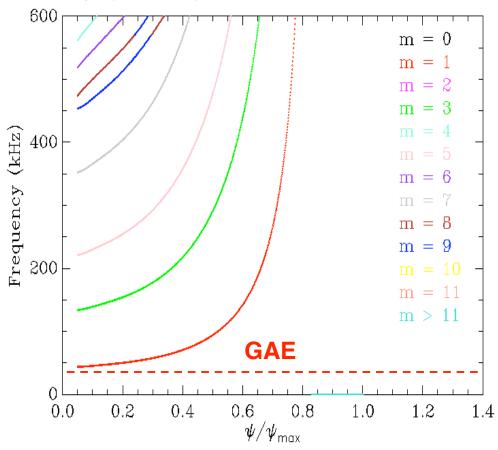
2. unstable when:  $\square^*_{dia} > \square_{Alfven}$  where  $\square^*_{dia}$  is the diamagnetic drift frequency

energetic ions or electrons can drive instability

## HSX : *Quasi-Helically Symmetric* (QHS) configuration Normal mode Alfvén continua: n = 1 mode family

#### Quasi-Helical Symmetry: Helical axis of symmetry, no toroidal curvature

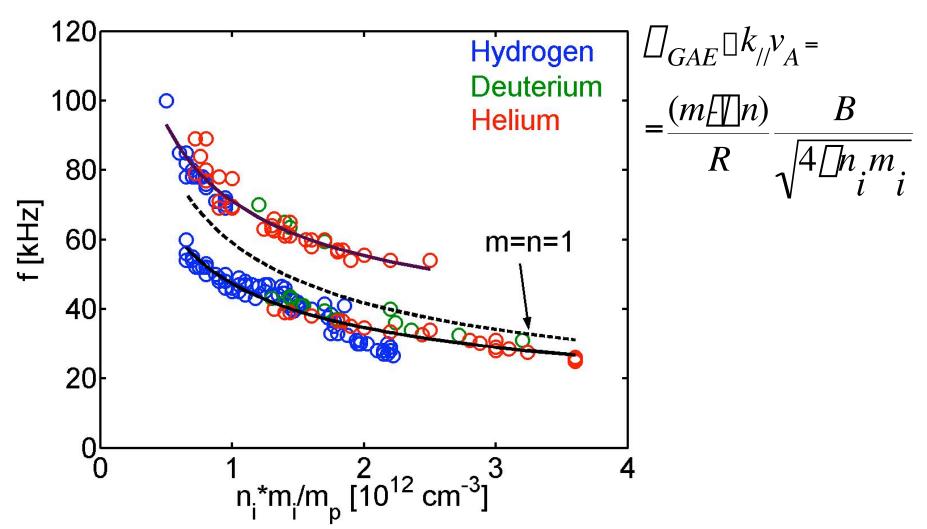
- GAE Gap: B=0.5 T 0 - 50 kHz for m=1,n=1 $n_e(0)=1.8 \times 10^{12} \text{ cm}^{-3}$
- Only minor changes for mirror configuration



B=0.5 T

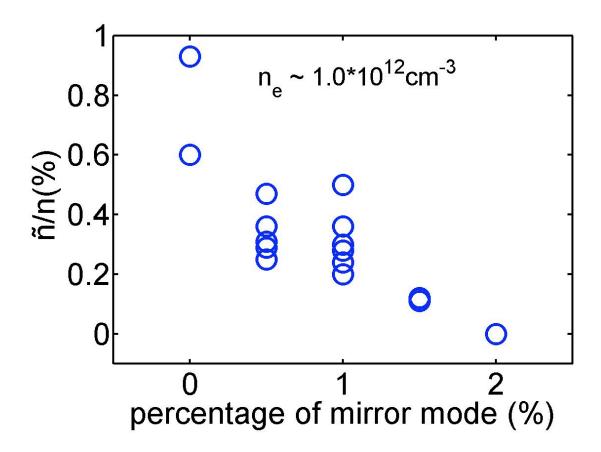
STELLGAP code (D. Spong)

## Mode frequency scaling with ion mass density



- frequency and mass density scaling consistent with Alfvenic mode
- If iota is lowered < 1, GAE gap disappears and mode not observed

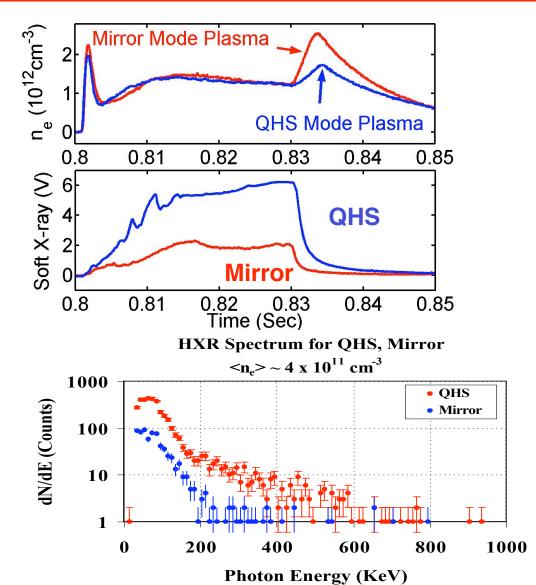
# Density fluctuations decrease with introduction of symmetry breaking (toroidal mirror) term



Fluctuation no longer observed for Mirror perturbation >2% (conventional stellarator configuration: ~10% mirror perturbation)

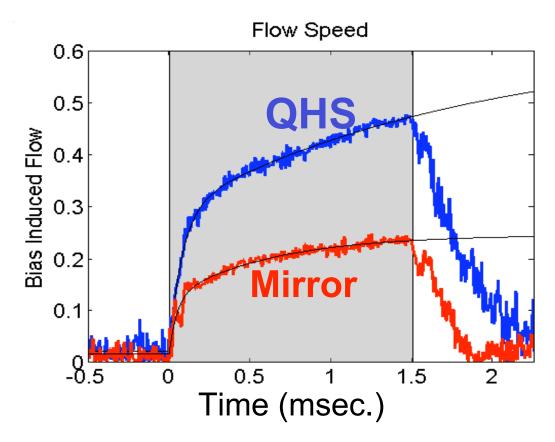
#### Soft X-ray, Hard X-ray Emission for QHS and Mirror

- Soft X-ray (600 eV-6 keV) emission
   QHS>>Mirror
- Hard X-ray flux: QHS>>Mirror decay time longer
- fast particles are better confined in QHS
- .□<sub>Dh</sub>=□<sub>GAE</sub>:
  5-10 keV particles



- fast particles (trapped electrons) are better confined for QHS
- provide drive for Alfvenic modes

# Result: QHS Flows Damp More Slowly, and, Go Faster For Less Drive *Viscous Damping is Reduced for QHS*



# QHS: 8 A of electrode current

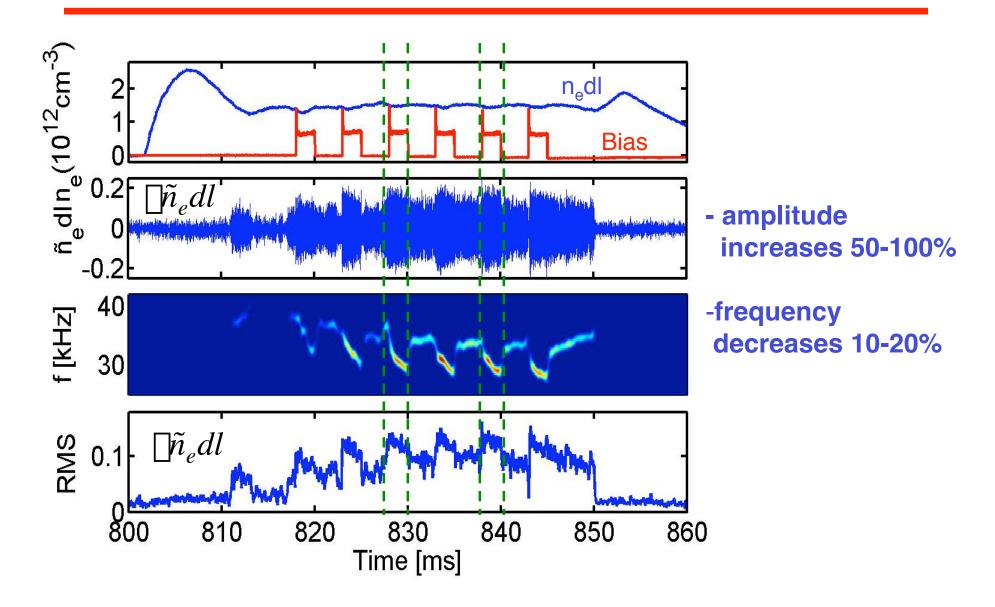
$$M = \frac{V_f}{\sqrt{T_e/m_i}} \square 0.5$$

Mirror: 10 A of electrode current

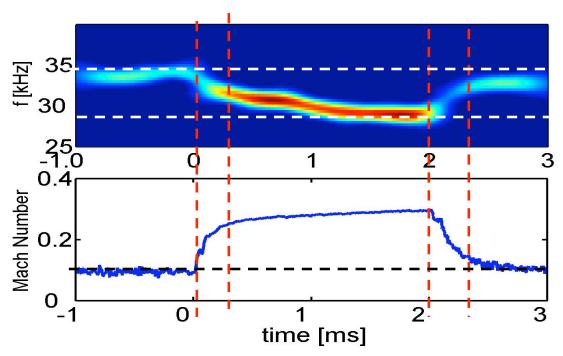
other parameters ( $n_e$ =1x10<sup>12</sup>cm<sup>-3</sup>,  $n_n$  | 1x10<sup>10</sup>cm<sup>-3</sup>  $T_i$ |25eV, B=0.5T,  $P_{ECH}$ =50 kW) held constant.

S.P. Gerhardt et al., PRL 94,015002(2005)

#### QHS: + biasing increases amplitude and decreases frequency



# Alfvenic mode frequency shift can be used to measure *core* flow dynamics



During biasing: n<sub>e</sub> and B do not change so V<sub>A</sub> is constant

Ambient plasma potential is (+)

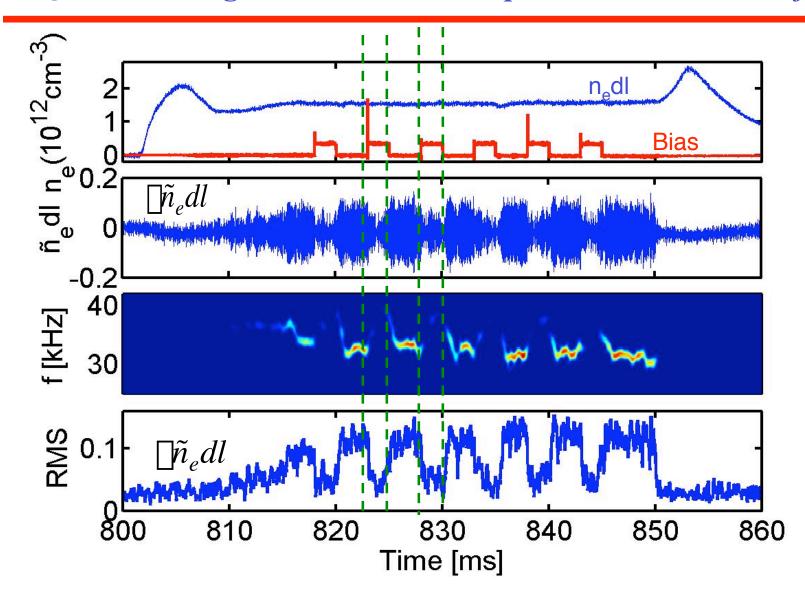
**ExB** flow in ion drift direction

Alfvenic mode propagates in electron diamagnetic drift direction?

$$f_{lab} = f_{GAE} + f_{Doppler} = f_{GAE} + \frac{1}{2 \square} k \bullet v_{Doppler} = f_{GAE} + \frac{m}{r} \frac{E_r}{2 \square B_o}$$

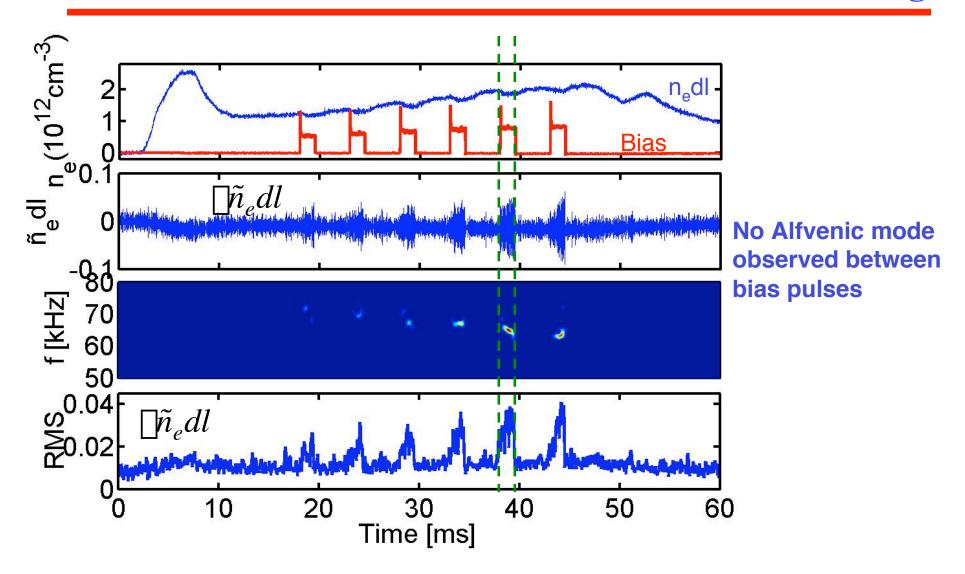
$$[]f_{lab} \ [] \ 5 \ [] \ 10 \ kHz; \ [] \ []E_r \ [] \ 5 \ [] \ 10 \ V/cm$$
 for plasma core with bias

QHS: - biasing decreases mode amplitude and increases frequency



- Biasing against direction of ambient flow

## Mirror Mode: Alfvenic Mode observed with + biasing



Mirror Mode: mode observed w/bias in direction of ambient flows no mode observed for opposite bias

#### **Evidence for Alfvenic mode in HSX**

- 1. Calculations of Alfven Wave Continuum by 3-D STELLGAP code shows the possibility of GAE mode in HSX
- 2. Measure a coherent fluctuation global mode [m=odd (1?)] with frequency and ion mass density scaling is consistent with Alfvénic mode (B scaling unknown).
- 3. Measurements suggest that the fluctuation is most likely driven by nonthermal electrons
- 4. Alfvenic Mode is only observed for QHS configuration, not for Mirror Configuration (2%)
- 5. Biasing: □f<sub>lab</sub> may provide information on core E<sub>r</sub> and flow dynamics!
   How do flows affect to Alfvenic mode growth rate?

Mode amplitude can be controlled by (1) flows and (2) configuration